

Neutrinos

Effective solar fusion process



Solar neutrino puzzle.

Calculated experimental detection rate

$$1 \text{ SNU} = 10^{-36} \text{ events / (target atom) / sec}$$

Observed in Davis chlorine $^{37}\text{Cl} = 3 \text{ SNU}$
 Calculated $6 \text{ SNU} = 2 \times (\text{observed})$

SNO detector (Sudbury Canada)

measured missing rate by observing

all 3 Flavors of neutrino. 2015 Nobel Prize

3 families of quarks and leptons

up	charm	top] quarks
down	strange	bottom	
ν_e	ν_μ	ν_τ] leptons
e	μ	τ	

3 flavors of quarks are produced with associated lepton, interacting with weak force carrier W :

$$W^- \rightarrow \begin{cases} \bar{\nu}_e + e^- \\ \bar{\nu}_\mu + \mu^- \\ \bar{\nu}_\tau + \tau^- \end{cases}$$

detected by production of corresponding charged lepton

$$\nu_e + W^- \rightarrow e^-$$

$$\nu_\mu + W^- \rightarrow \mu^-$$

$$\nu_\tau + W^- \rightarrow \tau^-$$

Solar neutrino puzzle solved by flavor oscillation

Sun $e^+ 2e^-$ \rightarrow earth ν_e

in fact $\nu_e \rightarrow \nu_\mu$ observe μ^+
 $\nu_e \rightarrow \nu_\tau$ observe τ^+

ν mixing

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{bmatrix} 3 \times 3 \\ \text{mixing} \end{bmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

flavor mass

3x3 mixing parameterized by
3 angles + phase

Simple two flavor mixing

$$\begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \begin{pmatrix} \cos\theta & \sin\theta \\ \sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} |\nu_1\rangle \\ |\nu_2\rangle \end{pmatrix}$$

ν are ultra relativistic

$$E = \sqrt{p^2 + m^2} \cong p + \frac{m^2}{2p} = pc +$$

$$e^{-iEt/\hbar} = e^{-ip/\hbar} e^{-im^2 t / 2p\hbar}$$

$$|\nu_e\rangle = e^{-ip/\hbar} \left[e^{-im_1^2 t / 2p\hbar} \cos\theta |\nu_1\rangle + e^{-im_2^2 t / 2p\hbar} \sin\theta |\nu_2\rangle \right]$$

ignore overall phase

$$|\nu_x(t)\rangle = \cos\theta |\nu_1\rangle + e^{-i\phi_{21}} \sin\theta |\nu_2\rangle$$

$$\phi_{21} \equiv \frac{m_2^2 - m_1^2}{2p\hbar} t \stackrel{p \approx E}{=} \left(\frac{\Delta m_{21}^2}{2E} \right) t$$

then at time t , amplitude to observe flavor state $|\nu_\mu\rangle$

$$|\nu_\mu\rangle = -\sin\theta |\nu_1\rangle + \cos\theta |\nu_2\rangle$$

$$\langle \nu_\mu | \nu_e \rangle = (-s, c) \begin{pmatrix} c \\ s e^{-i\phi_{21}} \end{pmatrix} = -sc(1 - e^{-i\phi_{21}})$$

$$= -2i \sin\theta \cos\theta e^{-i\phi_{21}/2} \sin\left(\frac{\phi_{21}}{2}\right)$$

$$= -i \sin(2\theta) e^{-i\phi_{21}/2} \sin\left(\frac{\phi_{21}}{2}\right)$$

$$P(\nu_e \rightarrow \nu_\mu) = |\langle \nu_\mu | \nu_e(t) \rangle|^2 = \sin^2(2\theta) \sin^2\left(\frac{\phi_{21}}{2}\right)$$

with $L = ct$

$$\frac{\Delta\phi_{21}}{2} = \frac{\Delta m_{21}^2 c^2 L}{4\hbar c E_\nu} = 1.27 \frac{\Delta m^2 [\text{eV}^2] L [\text{km}]}{E_\nu [\text{GeV}]}$$

for example $\Delta m^2 = 3 \times 10^{-3} \text{eV}^2$, $E_\nu = 1 \text{GeV}$

max mixing: $\sin^2\left(\frac{\phi_{21}}{2}\right) = 1$ $\frac{\phi_{21}}{2} = \frac{\pi}{2}$ @ $L = 400 \text{km}$

ν mass limits

Katrin exp. end-point of tritium β decay

$$m_{\nu_e} < 0.8 \text{eV} @ 90\% \text{CL}$$

$\langle \nu \rangle$ mass from cosmological observations
(CMB, baryon acoustic oscillations, ...)

$$\sum m_\nu < 0.09 \text{eV} @ 95\% \text{CL}$$

Only 3 light ($m_\nu < \frac{m_Z}{2} \approx 45 \text{ GeV}$)
 flavors from Z width $\Gamma_Z \approx 500 \text{ MeV}$

measured from $Z \rightarrow$ hadrons resonance

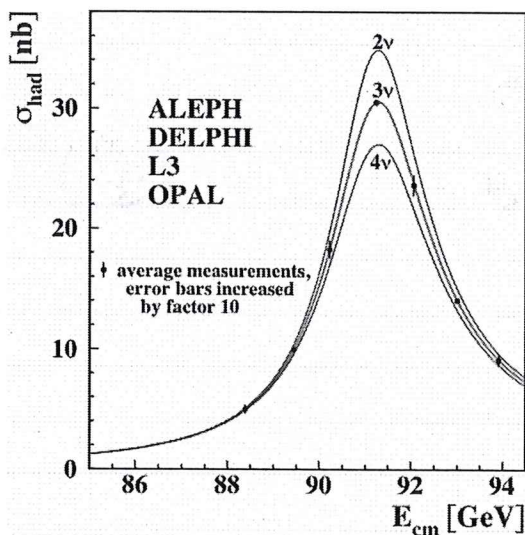


Fig. 1. Measurement of the hadron production cross-section as a function of the LEP centre-of-mass energy around the Z-boson resonance. Combined results from the four LEP experiments are presented. Curves represent the predictions for two, three and four neutrino species. To further convey the high sensitivity of the measurement, uncertainties are magnified tenfold.¹⁴

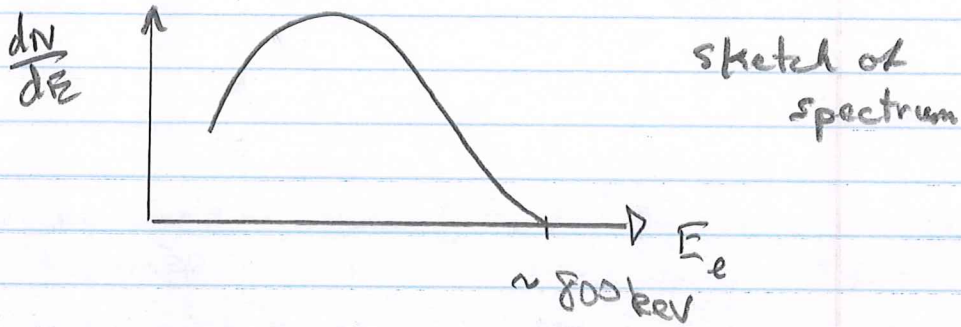
$$\Gamma_Z \approx \Gamma_{\text{had}} + \Gamma_{\text{ee}} + \Gamma_{\mu\mu} + \Gamma_{\tau\tau} + N_\nu \Gamma_{\nu\nu}$$

$$\underline{N_\nu = 3}$$

β decay

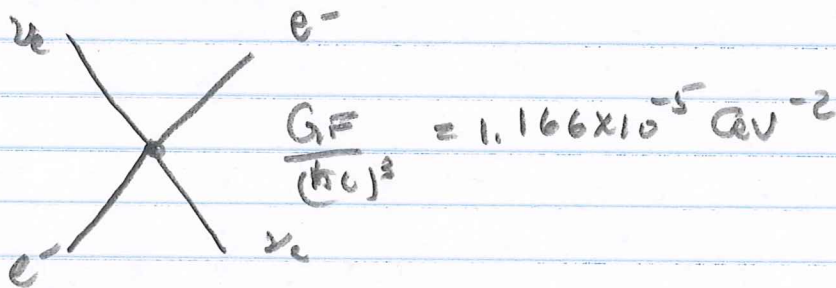
for example $n \rightarrow p + e^- + \bar{\nu}_e$

free neutron $\tau \approx 900s \approx 14$ minutes



E, \vec{p} conservation \Rightarrow unobserved ν from decay

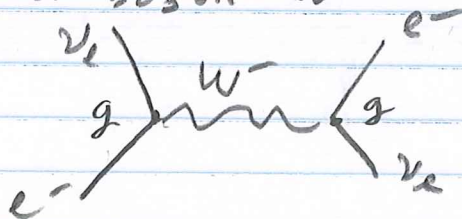
\propto Fermi contact interaction



$$\sigma = \frac{4}{\pi} G_F^2 (\hbar c k)^2 < \frac{\pi}{2k^2} \quad \text{S-wave unitarity}$$

$$\text{gives } (\hbar c k)^2 = \frac{\pi^2}{8G_F^2} \quad \hbar c k = 300 \text{ GeV limit}$$

Solution is to introduce intermediate vector boson W



V-A law

contact interaction is current current,
decay amplitude (matrix element)

$$\mathcal{M} = \frac{4G}{\sqrt{2}} J^\mu J_\mu^\dagger$$

where current is of the form

$$J_\ell^\mu = \bar{U}_\ell \Theta U_\ell \quad \ell = \text{lepton } e, \mu, \tau$$

five possible Lorentz invariant operators,

$$\Theta = 1, \gamma^5, \gamma^\mu, \gamma^5 \gamma^\mu, \sigma^{\mu\nu}$$

scalar, pseudoscalar, vector
axial-vector, tensor

$$\gamma^5 \equiv i \gamma^0 \gamma^1 \gamma^2 \gamma^3$$

$$\sigma^{\mu\nu} = \frac{i}{2} (\gamma^\mu \gamma^\nu - \gamma^\nu \gamma^\mu)$$

by experiment leptonic current is
pure V-A, parity violating:

$$J_\ell^\mu = \bar{U}_\ell \gamma^\mu \frac{1}{2} (1 - \gamma^5) U_\ell$$

$$(J_\ell^\dagger)_\mu = \bar{U}_\ell \gamma_\mu \frac{1}{2} (1 - \gamma^5) U_\ell$$

ν left handed, $\bar{\nu}$ right handed

Muon decay $\mu^- \rightarrow \nu_\mu + e^- + \bar{\nu}_e$

Putting in W ,

$$M = \left(\frac{g}{\sqrt{2}}\right) \bar{\psi}_\mu \left(\frac{1}{m_W^2 - \partial^2}\right) \frac{g}{\sqrt{2}} (\psi_e^+)$$

↳ Lorentz index

$$\xrightarrow{\partial^2=0} \frac{g^2}{2} \frac{1}{m_W^2} \bar{\psi}_\mu (\psi_e^+)$$

then $\boxed{\frac{4G_F}{\sqrt{2}} = \frac{g^2}{2} \frac{1}{m_W^2}}$

from muon lifetime, $E_0 = E_e + E_\nu$

$$\Gamma = \frac{1}{\tau} = \frac{G_F^2 m_\mu^5}{192 \pi^3} = 2.2 \mu\text{s}$$

$$\alpha_W = \frac{g^2}{4\pi} = \frac{1}{29} > \alpha_{EM}$$

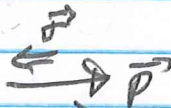
$$m_W = 80 \text{ GeV}$$

weak force is weak because m_W is heavy

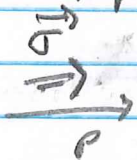
"handedness" chirality

In relativistic (massless) limit (chirality = helicity)

ψ left chiral



$\bar{\psi}$ right chiral



$$\underline{\gamma^5} : \{ \gamma^5, \gamma^{\mu\nu} \} = 0$$

$$(\gamma^5)^2 = I ; (\gamma^5)^\dagger = \gamma^5$$

in Pauli-Dirac representation

$$\gamma^0 = \begin{pmatrix} I & 0 \\ 0 & -I \end{pmatrix} \text{ diagonal mass}$$

$$\vec{\gamma} = \begin{pmatrix} 0 & \vec{\sigma} \\ -\vec{\sigma} & 0 \end{pmatrix} \quad \gamma^5 = \begin{pmatrix} 0 & I \\ I & 0 \end{pmatrix}$$

$$\text{so } \gamma^\mu \left(\frac{1-\gamma^5}{2} \right) = \left(\frac{1+\gamma^5}{2} \right) \gamma^\mu$$

$$\text{and } \left(\frac{1-\gamma^5}{2} \right)^2 = \frac{1-\gamma^5}{2}$$

$$\text{then } \gamma^\mu \left(\frac{1-\gamma^5}{2} \right) = \left(\frac{1+\gamma^5}{2} \right) \gamma^\mu \left(\frac{1-\gamma^5}{2} \right)$$

$$\text{defining } U_L = \left(\frac{1-\gamma^5}{2} \right) U$$

$$\text{so } \bar{U} = \bar{U} \left(\frac{1+\gamma^5}{2} \right)$$

and current is left-chiral, *chiral*

$$J_A^\mu = \bar{U}_L \gamma^\mu \frac{1}{2} (1-\gamma^5) U_L$$

$$= \bar{U}_L \gamma^\mu U_L$$

W^- couples to left-chiral current

parity violation \Rightarrow left-right symmetry breaking

Weyl equations

Dirac matrices in Pauli-Dirac representation diagonalize \not{p} (mass).
Weyl representation diagonalizes chirality,
more appropriate for weak interactions.

$$\gamma^5 = \begin{pmatrix} -\mathbb{I} & 0 \\ 0 & \mathbb{I} \end{pmatrix} \quad \gamma^0 = \begin{pmatrix} 0 & \mathbb{I} \\ \mathbb{I} & 0 \end{pmatrix}$$

$$\vec{\gamma} = \begin{pmatrix} 0 & \vec{\sigma} \\ -\vec{\sigma} & 0 \end{pmatrix}$$

then Dirac equation with plane wave solution

$$\psi = U(\vec{p}) e^{i\vec{p}\cdot\vec{x}} \quad u$$

$$\begin{pmatrix} -\vec{\sigma}\cdot\vec{p} & m \\ m & \vec{\sigma}\cdot\vec{p} \end{pmatrix} \begin{pmatrix} \chi \\ \phi \end{pmatrix} = E \begin{pmatrix} \chi \\ \phi \end{pmatrix}$$

we see that mass couples L, R chiral wave functions.

Chiral projection operators,

$$P_L = \frac{1}{2}(1 - \gamma^5) = \begin{pmatrix} \mathbb{I} & 0 \\ 0 & 0 \end{pmatrix}$$

$$P_R = \frac{1}{2}(1 + \gamma^5) = \begin{pmatrix} 0 & 0 \\ 0 & \mathbb{I} \end{pmatrix}$$

$$\psi_L = P_L \psi = \begin{pmatrix} \chi \\ 0 \end{pmatrix} \quad \psi_R = P_R \psi = \begin{pmatrix} 0 \\ \phi \end{pmatrix}$$

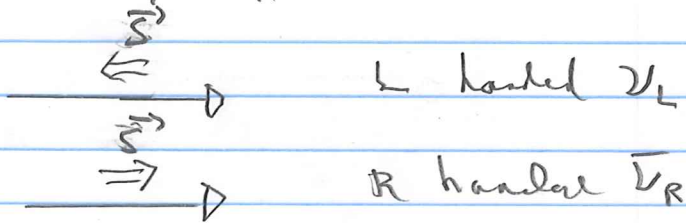
4 comp. ψ is direct sum of spinor representations of Lorentz group.

Ultra relativistic limit ($m \ll 0$)

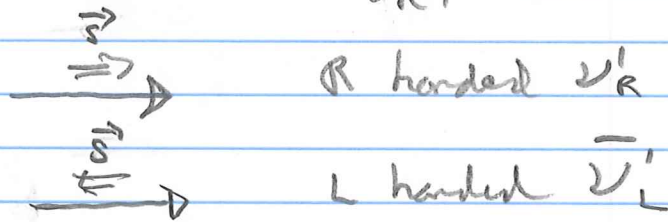
$$-\vec{\sigma} \cdot \vec{p} \chi = E \chi \quad E > 0$$

$$\vec{\sigma} \cdot \vec{p} \phi = E \phi \quad E < 0$$

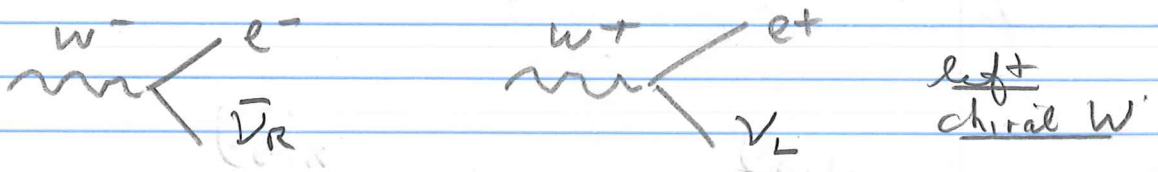
χ describes $\nu_L, \bar{\nu}_R$ when



when re-interpreted as $E > 0$ solution, bottom equation describes $\nu'_R, \bar{\nu}'_L$

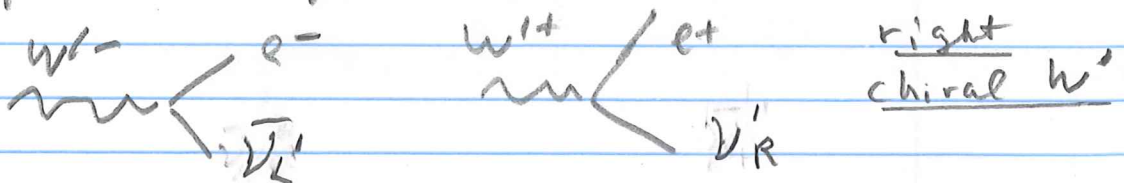


Weak interactions couple to top state (ν) as



left chiral w

Bottom states are sterile with respect to weak interactions w, z . They would couple to a hypothetical w'



right chiral w'

W' direct search (Atlas @ LHC)

$$m_{W'} \gtrsim 1500 \text{ GeV (95\% CL)}$$

Returning to Pauli-Dirac rep. to look at chirality,

$$\gamma^5 = \begin{pmatrix} 0 & I \\ I & 0 \end{pmatrix}$$

Dirac

spinor

$$U^{S'}(\vec{p}) = \sqrt{E+m} \begin{pmatrix} \chi^S \\ \frac{\vec{\sigma} \cdot \vec{p}}{E+m} \chi^S \end{pmatrix}$$

$$\xrightarrow{E \gg m}$$

$$\sqrt{E} \begin{pmatrix} \chi^S \\ \vec{\sigma} \cdot \hat{p} \chi^S \end{pmatrix}$$

$$\hat{p} = \vec{p}/|\vec{p}| \\ = \vec{p}/E$$

so

$$\gamma^5 U^{S'}(\vec{p}) = \sqrt{E} \begin{pmatrix} \vec{\sigma} \cdot \hat{p} \chi^S \\ \chi^S \end{pmatrix}$$

since $(\vec{\sigma} \cdot \hat{p})^2 = I$

$$= \sqrt{E} \begin{pmatrix} \vec{\sigma} \cdot \hat{p} \chi^S \\ (\vec{\sigma} \cdot \hat{p})^2 \chi^S \end{pmatrix}$$

$$= \begin{pmatrix} \vec{\sigma} \cdot \hat{p} & 0 \\ 0 & \vec{\sigma} \cdot \hat{p} \end{pmatrix} U^{S'}$$

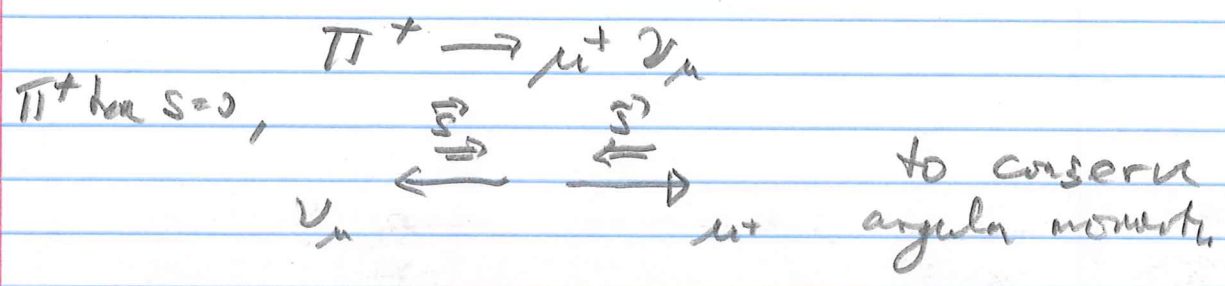
$$= \vec{\sigma} \cdot \hat{p} U^{S'}$$

helicity operator. in ultrarelativistic limit,

chirality = helicity

Note in Weyl rep. $\vec{\Sigma} \cdot \vec{p}$ is the same.

For non-relativistic particles, for example μ^+ produced in weak decay



μ^+ produced as right-chiral from V-A law, probability to be in left helicity state is.

Prob_L = $\frac{1}{2}(1 - \frac{v}{c})$ $\xrightarrow{v/c=1}$ 0

and right helicity $\xrightarrow{v/c=1}$

Prob_R = $\frac{1}{2}(1 + \frac{v}{c})$ $\xrightarrow{v/c=1}$ 1

Since angular momentum forces μ^+ to be in left helicity state, decay is suppressed by factor $(1 - v/c)$.

e^+ in decay $\pi^+ \rightarrow e^+ \nu_e$ is more relativistic, so even more suppressed.

decay rate $m_\pi^+ \rightarrow l^+ \nu_l$ $l = e, \mu$

$\Gamma \propto (\text{phase space}) \times (1 - \frac{v}{c})$

Relativistic Kinematics

$$\begin{array}{c} \longleftarrow \quad \longrightarrow \\ p_z \quad p_x \end{array} \quad |\vec{p}_z| = |\vec{p}_x| = p$$

$$E_0 = m_\pi = p + \sqrt{p^2 + m^2} \Rightarrow p = \frac{m_\pi^2 - m^2}{2m_\pi}$$

phase space $p^2 \frac{dp}{dE_0}$

$$\frac{dE_0}{dp} = 1 + \frac{p}{\sqrt{p^2 + m^2}} = 1 + \frac{p}{m_\pi - p} = \frac{m_\pi}{m_\pi - p}$$

$$\begin{aligned} \frac{dp}{dE_0} &= \left(1 - \frac{p}{m_\pi}\right) = 1 - \frac{1}{2m_\pi^2} (m_\pi^2 - m^2) \\ &= \frac{1}{2} \left(\frac{m_\pi^2 + m^2}{m_\pi^2} \right) \end{aligned}$$

$$\text{then } p^2 \frac{dp}{dE_0} = \frac{1}{4m_\pi^4} (m_\pi^2 + m^2) (m_\pi^2 - m^2)^2$$

$$E^2 = p^2 + m^2 = \frac{(m_\pi^2 + m^2)^2}{4m_\pi^2}$$

$$\text{so } \frac{p}{E} = \frac{v}{c} = \frac{(m_\pi^2 - m^2)}{(m_\pi^2 + m^2)}$$

$$\text{sim } 1 - \frac{v}{c} = \frac{2m^2}{m_\pi^2 + m^2}$$

then decay rate

$$\Gamma \propto p^2 \frac{dp}{dE} \left(1 - \frac{v}{c}\right)$$

$$= \frac{1}{4m_\pi^4} (m_\pi^2 + m^2) (m_\pi^2 - m^2)^2 \left(\frac{2m^2}{m_\pi^2 + m^2}\right)$$

$$= \frac{1}{4m_\pi^2} 2m^2 (m_\pi^2 - m^2)^2$$

$$\boxed{\Gamma \propto \frac{m^2}{2} \left(1 - \frac{m^2}{m_\pi^2}\right)^2}$$

$$m_\pi = 135 \text{ MeV}, \quad m_\mu = 106 \text{ MeV}, \quad m_e = 0.5 \text{ MeV}$$

so neglect m_e ,

$$\frac{\Gamma(\pi^+ \rightarrow e^+ \nu_e)}{\Gamma(\pi^+ \rightarrow \mu^+ \nu_\mu)} = \frac{m_e^2}{m_\mu^2} \frac{1}{\left(1 - \frac{m_\mu^2}{m_\pi^2}\right)^2}$$

$$= 1.3 \times 10^{-4}$$

called helicity suppression

Majorana ν ?

electron has 4 states

e^- (spin up, down), e^+ (spin up, down)

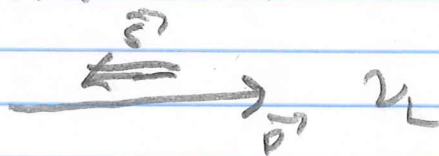
or in terms of chirality

$$P_R + P_L = \frac{1}{2}(1 + \gamma_5) + \frac{1}{2}(1 - \gamma_5) = I$$

particle	e_L^-	e_R^-	:	ν_L	ν_R'
antiparticle	e_R^+	e_L^+	:	$\bar{\nu}_R$	$\bar{\nu}_L'$
(denoted by charge or over-bar)	electron			neutrino	

as far as we know ν_R' , $\bar{\nu}_L'$ do not exist.
(still looking)

then neutrino states are



what all we know is we see

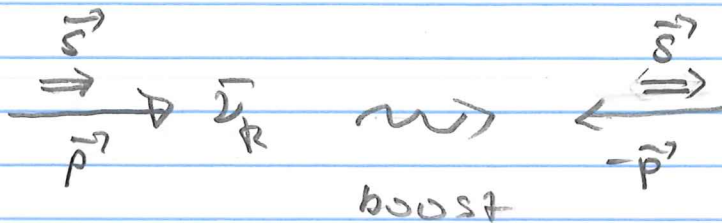
$$l^- \bar{\nu}_R, e^+ \nu_L$$

right handed
anti-neutrino

Left handed
neutrino

We can Lorentz boost to

frame where $\vec{p} \rightarrow -\vec{p}$, spin does not change



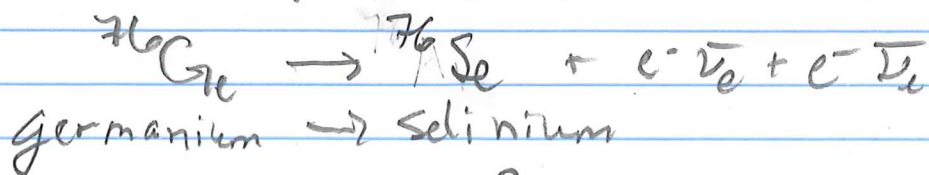
in boosted frame $\bar{\nu}_R$ is just like ν_L .

Could it be the same state? Called Majorana ν .

[Could anti-neutrino $\bar{\nu}_R$ be the same as neutrino ν_L with just opposite helicity?]

Since ν are ultra-relativistic we can never do this experiment.
Only way to know is to look for neutrinoless double- β decay.

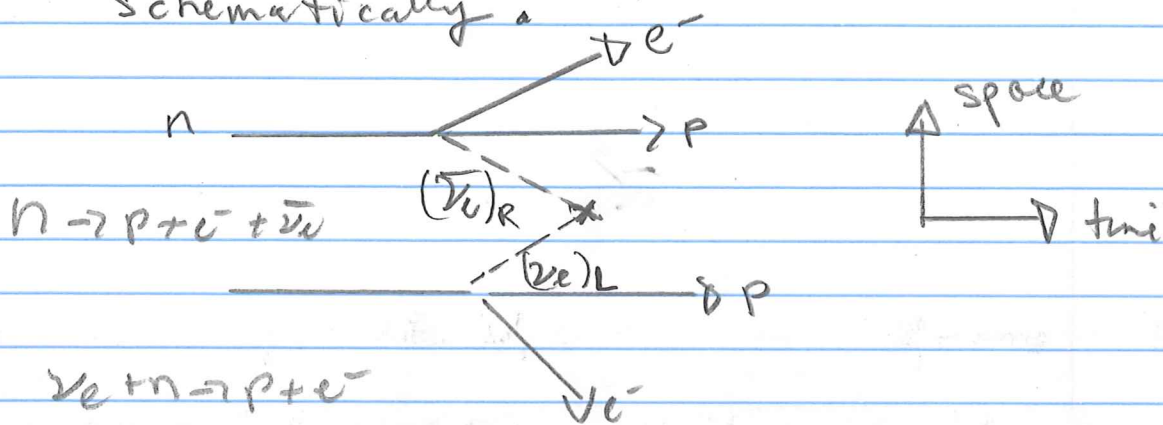
Some nuclei, e.g. ^{76}Ge are Pauli-blocked from beta decay, can undergo double beta decay



$$T_{1/2} \approx 2 \times 10^{21} \text{ years.}$$

Since we can think of anti particles as particle moving backwards in time,

schematically:



then $n \rightarrow p p e^- e^-$ no neutrinos!

can only happen if $\bar{\nu}_e, \nu_e$ are particle, antiparticle (Majorana ν).

Expected half-life $\sim 10^{25}$ years.

If ν is Majorana, then cannot have Dirac mass by interacting with Higgs like other fermion unless

ν' exists or

Some new physics creates Majorana mass term. Also, Majorana ν introduces 2 new CP phases in mixing matrix. Not observable in oscillation experiments. Possible solution to problem of baryogenesis (matter/antimatter) asymmetry puzzle - universe is all matter.