

Lecture #2I Cross sections & Decay Rates

primary observables in nuclear & particle physics

① decay rate $\Gamma = \text{constant prob. / time}$

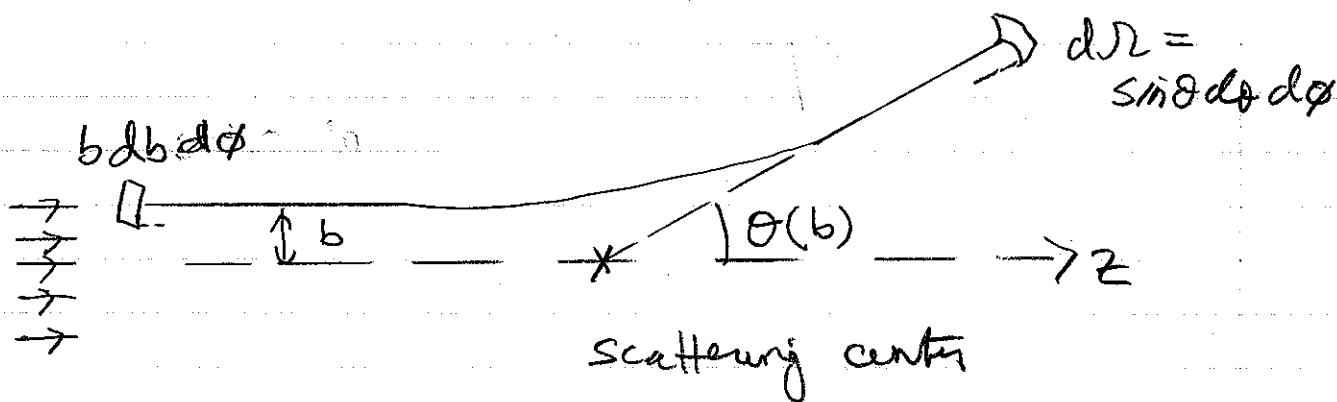
$$dN = -\Gamma N dt$$

$$\langle t \rangle = 1/\Gamma \equiv \tau \quad \text{"lifetime"}$$

② cross section flux normalized scattering rate

Classical: particles follow trajectory.

Scattering angle θ depends on impact parameter b



flux $F \equiv \# \text{ incident particles / area / time}$

$\phi \equiv \text{azimuthal angle}$

all particles thru $b db d\phi \rightarrow d\Omega$

$$\Delta\sigma = \frac{(\text{rate @ } \theta, \phi)}{F} \cdot \Delta\Omega = \frac{F \cdot b \Delta b \Delta\phi}{F}$$

$$\frac{\Delta\sigma}{\Delta\Omega} = \frac{b \Delta b \Delta\phi}{\sin\theta \Delta\theta \Delta\phi} = \frac{b}{\sin\theta} \frac{\Delta b}{\Delta\theta}$$

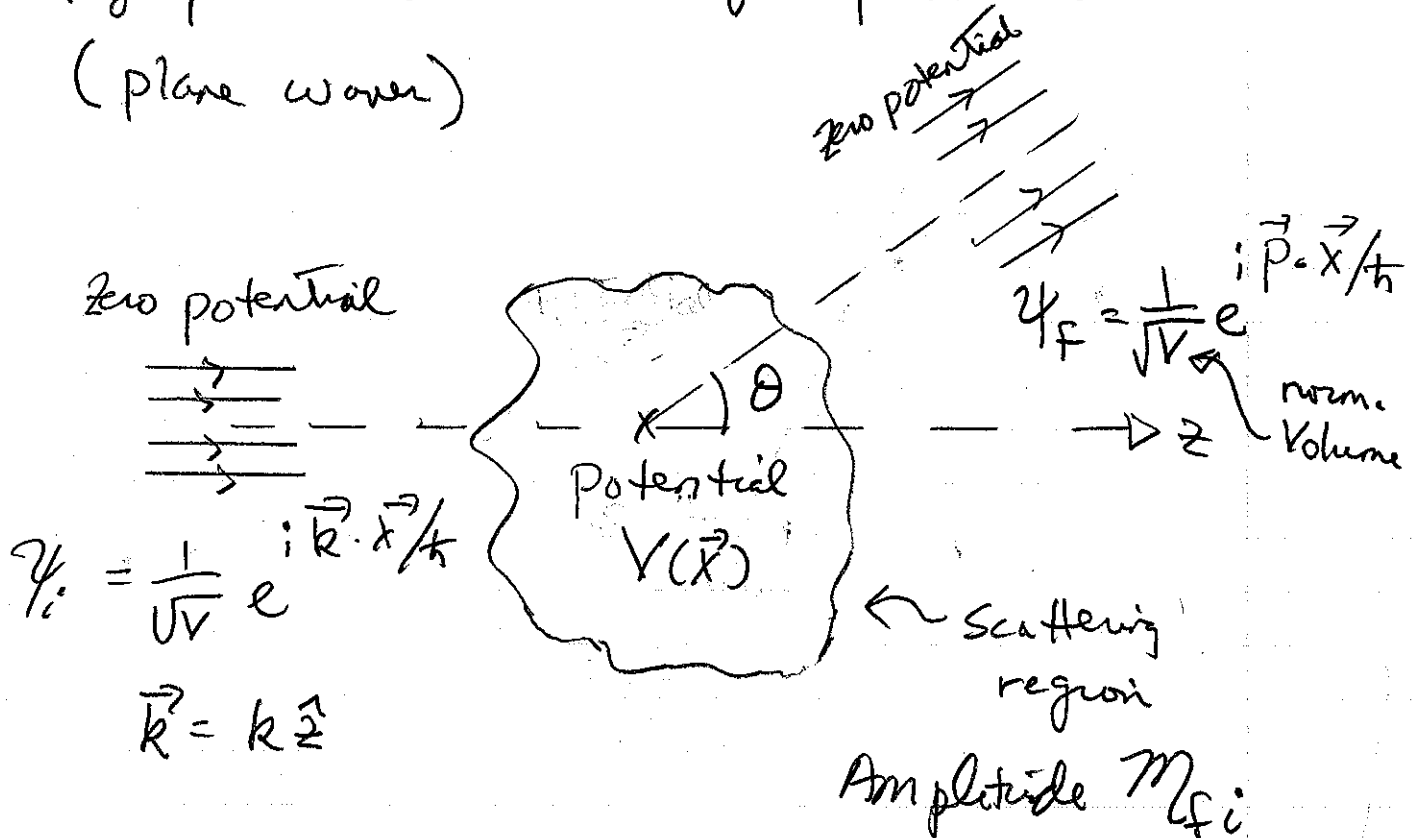
$$\frac{d\sigma}{d\Omega} = \frac{b}{\sin\theta} \left| \frac{db}{d\theta} \right|$$

③ Non-Relativistic Quantum scattering

No trajectories; amplitude to scatter

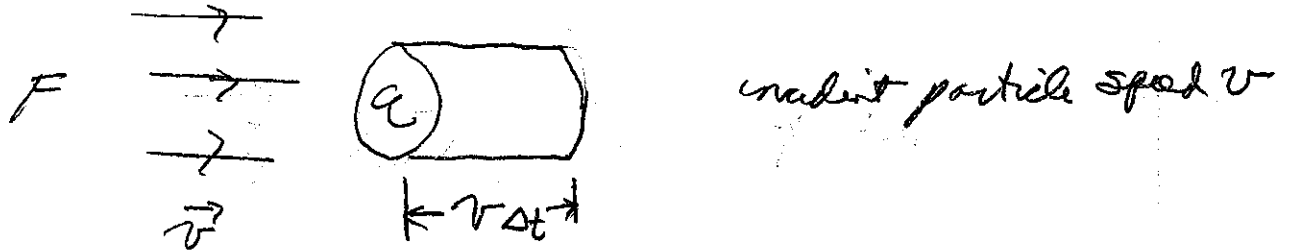
Asymptotic states are free particles

(plane waves)



$$dR = \frac{dR}{F}$$

Q.M. flux includes incident wave function normalization of 1 particle/volume.



$$F = \left(\frac{1}{V} \right) \frac{v a t a}{a t} = \frac{v}{V}$$

④ Fermi's golden rule transition rate dR ,

$$dR = \frac{2\pi}{\hbar} |M_{fi}|^2 \rho(E)$$

Amplitude, dynamics \uparrow density of states, kinematics

$$\rho(E) = \frac{dN_S}{dE} = \frac{V}{(2\pi\hbar)^3} \frac{p^2 dp d\Omega}{dE} = \frac{V}{(2\pi\hbar)^3} \frac{p^2}{dE/dp} d\Omega$$

$$\frac{dE}{dp} = \frac{d}{dp} \left(\frac{p^2}{2m} \right) = \frac{p}{m} = v$$

note: relativistic case is the same,

$$\frac{dE}{dp} = \frac{d}{dp} \left(\sqrt{p^2 + m^2} \right) = \frac{p}{E} = \frac{mv}{m\gamma} = v$$

⑤ Scattering Amplitude

In first order perturbation theory:

$$m_{fi}^{(1)} = \int \psi_f^* V(\vec{x}) \psi_i d^3x$$

$$= \frac{1}{V} \int e^{i\vec{q}\cdot\vec{x}/\hbar} V(\vec{x}) d^3x$$

where $\vec{q} \equiv \vec{k} - \vec{p}$ is momentum transfer.

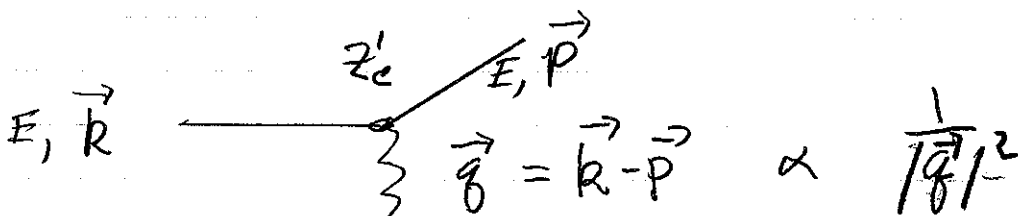
m is just Fourier Transform of potential

For $V(\vec{x}) = \frac{zz'e^2}{|\vec{x}|}$,

$$m_{fi}^{(1)} = \frac{zz'e^2}{V} \int e^{i\frac{\vec{q}\cdot\vec{x}}{\hbar}} \frac{d^3x}{|\vec{x}|} = \frac{zz'e^2}{V} \cdot \frac{4\pi\hbar^2}{|\vec{q}|^2}$$

Choose spherical coordinates with \vec{q} as z-axis for integration.

Diagram representation (Feynman-like)



Note: relativistic $\vec{q} = (0, \vec{q})$ exchange of virtual photon $\vec{q}\cdot\vec{q} \equiv t \neq 0$ "off mass shell"

⑥ Rutherford Cross Section

$$d\sigma = \left(\frac{V}{v}\right) \frac{2\pi}{\hbar} \left(\frac{ZZ'e^2}{V}\right)^2 \frac{(4\pi\hbar^2)^2}{g^4} \frac{V}{(2\pi\hbar)^3} \frac{p^2}{v} d\Omega$$

$$= 4 \frac{p^2}{v^2} \frac{1}{g^4} (ZZ'e^2)^2$$

non-relativistic $p/v = m$; $E = \frac{p^2}{2m}$

$$g^2 = |\vec{k} - \vec{p}|^2 = 2p^2(1 - \cos\theta) = 4p^2 \sin^2 \theta/2$$

$$\frac{d\sigma}{d\Omega} = \frac{1}{4} \left(\frac{m}{p^2}\right)^2 \frac{(ZZ'e^2)^2}{\sin^4(\theta/2)}$$

$$= \frac{1}{(4E)^2} \frac{(ZZ'e^2)^2}{\sin^4(\theta/2)}$$

no $\hbar \Rightarrow$ same as classical result.

⑦ Relativistic Cross Section

Relativistic scattering cross section is similar, written in terms of invariant amplitude M_{fi} :
 $1+2 \rightarrow 3 \dots n+2$ (n final particles)

$$d\sigma = \underbrace{\frac{1}{2E_1 2E_2}}_{\text{invariant flux}} \left(\frac{1}{v_1 + v_2}\right) |M_{fi}|^2 \underbrace{d\Phi_n}_{\text{invariant phase space}}$$

note: dot being area \perp beam is Lorentz Invariant.

where invariant flux:

$$\vec{k}_1 = E_1 (1, \vec{v}_1)$$

$$\vec{k}_2 = E_2 (1, \vec{v}_2)$$

$$\text{for } \vec{v}_1 \cdot \vec{v}_2 = +v_1 v_2, \quad |\vec{v}_1 + \vec{v}_2|^2 = [v_1^2 + v_2^2 + 2v_1 v_2]^{1/2}$$

$$= v_1 + v_2$$

$$2E_1 E_2 (v_1 + v_2) = 4 \sqrt{(\vec{k}_1 \cdot \vec{k}_2)^2 - m_1^2 m_2^2}$$

invariant phase space:

$$d\Phi_n = (2\pi)^4 \delta^4(\vec{k}_1 + \vec{k}_2 - \sum_{i=1}^n \vec{p}_i) \prod_{i=1}^n \frac{d^3 p_i}{2E_i (2\pi)^3}$$

Similarly, decay of particle w/ mass $M \rightarrow n$ particles
in rest frame

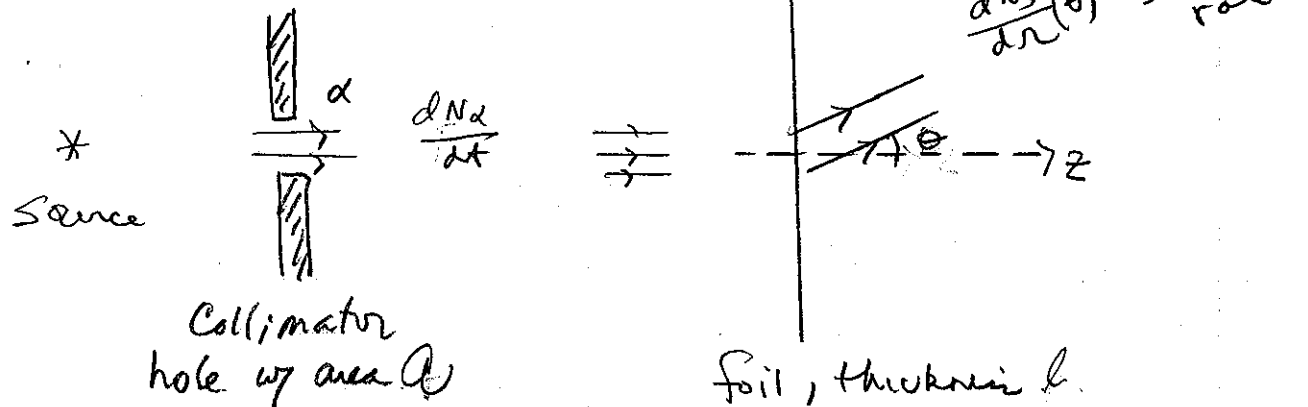
$$d\Gamma = \frac{1}{2M} |m|^2 d\Phi_n$$

Formulas are nice looking
after lots of messaging and moving
around of factor of $E, 2\pi$.

note: $\frac{d^3 p}{E}$ is invariant.

The Rutherford Experiment

Thin foil scattering



"thin" means ignore multiple scattering

$$F = \frac{dN_\alpha}{dt} \left(\frac{1}{a} \right) \times (\# \text{ targets})$$

(# α 's / time) thru hole

$$\# \text{ targets} = \frac{A_0 \rho}{M_A} (la)$$

#/vol

A_0 = Avogadro's number

$M_A = A \times 10^{-3} \text{ kg/mole} = A \times 1 \text{ gm/mole}$ molar mass

ρ = density gm/vol

$$F = \frac{dN_\alpha}{dt} \left(\frac{A_0 \rho}{M_A} \right) l$$

Measure cross section from observed scattering rate dN_s/dt :

$$\frac{d\Gamma_{\text{exp}}}{d\Omega}(\theta) = \frac{1}{F} \frac{dN_s}{dt}(\theta) = \left(\frac{m_A}{A_0 S R} \right) \left(\frac{dN_s(\theta)/dt}{dN_i/dt} \right)$$

Agreement with theoretical Rutherford formula

Verifies theoretical assumptions:

- ① mass assumed infinite

$$M_N \approx M_{\text{Atom}}$$

- ② size is "pointlike" $\ll 0.1 \text{ nm}$ atomic scale

"pointlike" means follows $\frac{d\sigma}{d\Omega} \propto \frac{1}{\sin^4 \theta/2} \Big|_{\theta=\pi} = 1$

Extended object σ would be expected to decrease faster w/ θ .

From Text by Roy & Nigam

8 NUCLEAR SIZE AND NUCLEAR SHAPES

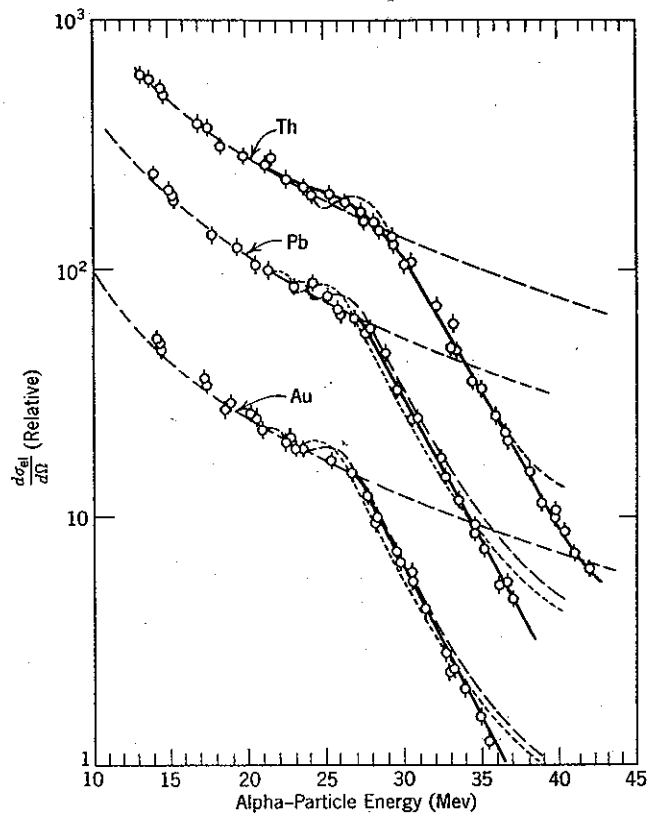


Fig. 2-1 The broad dashed curve gives the Coulomb cross section, and the solid curve represents the experimental data of Farwell and Wegner⁴ for Au, Pb, and Th. For Au, the finer theoretical curve corresponds to $R = 10.58 \times 10^{-13}$ cm and the coarser curve to $R = 10.3 \times 10^{-13}$ cm. For Pb, the finer curve corresponds to $R = 10.87 \times 10^{-13}$ cm and the coarser to $R = 10.42 \times 10^{-13}$ cm. For Th, the dashed curve corresponds to $R = 11.01 \times 10^{-13}$ cm (Eisberg and Porter¹).

--- Rutherford scattering

Distance scale probed by α :

Classical minimum distance r_{\min} :

$$E_{\alpha} = \frac{ZZ'e^2}{r_{\min}}$$

$$\frac{e^2}{\hbar c} = \alpha \approx \frac{1}{137} \quad \text{fine structure constant}$$

take $E_{\alpha} = 4 \text{ MeV}$ on ${}^{207}_{82}\text{Pb}$

$$r_{\min} = \frac{2(82)}{137} \frac{200 \text{ MeV}}{4 \text{ MeV}} = \frac{8200}{137} \text{ fm}$$

$$= 60 \text{ fm}$$

nuclear size $r_n \approx 1.2 \text{ fm } A^{1/3}$

$$= 1.2 (207)^{1/3} \text{ fm} = 7 \text{ fm}$$

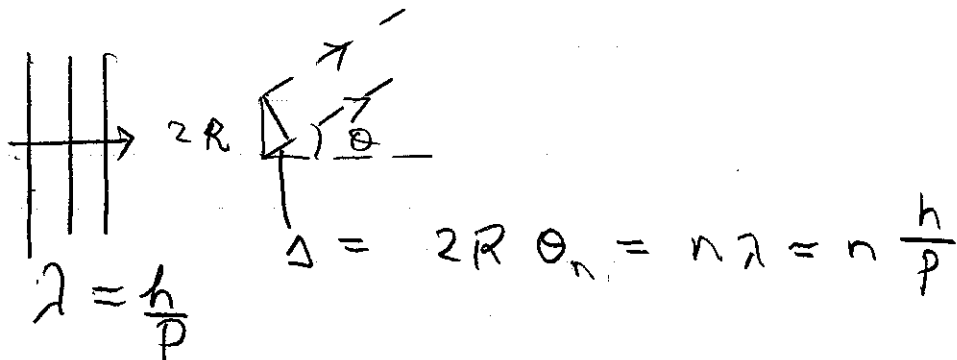
$$r_{\min} \gg r_{\text{pb}}$$

for $E = 25 \text{ MeV}$, $r_{\min} \approx 10 \text{ fm}$. Begin to probe nuclear size.

low energy

elastic π -nucleus scattering: $\frac{d\sigma}{d\Omega}(\theta)$

shows structure that can be interpreted as wave scattering by absorptive sphere:



note: first dip for π -Pb scattering is slightly above 50°

Mott cross section

Relativistic ($v \approx 1$) e^- elastic scattering
by spinless, point-like nucleus of mass
 M , charge Ze :

$$\frac{d\sigma^{\text{Mott}}}{d\Omega} = \frac{(Z\alpha\hbar c)^2}{4E_0^2 \sin^4 \frac{\theta}{2}} \left(\frac{E}{E_0} \right) \cos^2 \frac{\theta}{2}$$

\uparrow \uparrow
 He coil e^- spin
 $v=1$

ultra-relativistic
 $E^2 = p^2$

In ultra-relativistic limit, helicity is conserved.

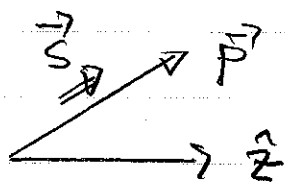
$$H = \left\langle \frac{\vec{S} \cdot \vec{p}}{|\vec{S}| |\vec{p}|} \right\rangle = \pm 1$$

$$\begin{array}{c} \uparrow \vec{S} \\ \rightarrow \vec{p} \end{array} \quad \text{"right handed"}$$

$$\begin{array}{c} \leftarrow \vec{S} \\ \rightarrow \vec{p} \end{array} \quad \text{"left handed"}$$

Incoming wave function with $\vec{p} = p\hat{z}$ has $L_z = 0$.
Conservation of angular momentum requires that

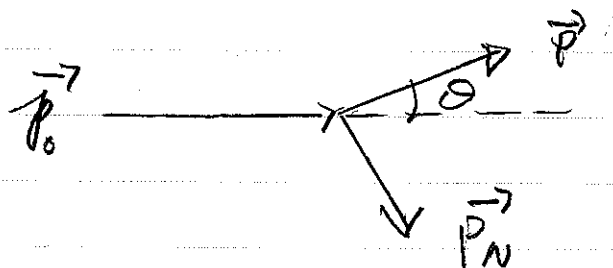
$\langle \vec{S} \cdot \hat{z} \rangle$ does not change



$$|S\rangle_{\vec{P}} = \left| \frac{1}{2}, \pm \frac{1}{2} \right\rangle_P$$

$$\text{Amplitude } \langle \frac{1}{2}, \pm \frac{1}{2} | \frac{1}{2}, \pm \frac{1}{2} \rangle_P = \cos \frac{\theta}{2}$$

Take nuclear recoil into account



$$\vec{P} = P_0(1, \hat{z}) \quad ; \quad \vec{P} = P(1, \hat{r}) \quad ; \quad \vec{P}_0 \cdot \vec{P} = 0 \quad ; \quad \vec{P} \cdot \vec{P} = 0$$

4 momentum transfer:

$$\begin{aligned} -q^2 &= -(\vec{P}_0 - \vec{P})^2 = 2PP_0(1 - \cos\theta) \\ &= 4E_0E \sin^2 \frac{\theta}{2} \end{aligned}$$

Energy conservation determines $E(\theta)$:

$$E_0 + m_N = E + E_N$$

$$(E_0 + m_N - E)^2 = E_N^2 = m_N^2 + |\vec{p}_N|^2$$

$$|\vec{p}_N|^2 = |\vec{p}_0 - \vec{p}|^2 = p_0^2 + p^2 - 2p_0 p \cos \theta$$

$$= E_0^2 + E^2 - 2E_0 E \cos \theta$$

$$= (E_0 - E)^2 + 4E_0 E \sin^2 \theta/2$$

$$(\cancel{E_0 - E})^2 + 2m_N(E_0 - E) = (\cancel{E_0 - E})^2 + 4E_0 E \sin^2 \theta/2$$

$$\left(\frac{1}{E} - \frac{1}{E_0} \right) = \frac{2 \sin^2 \theta/2}{m_N}$$

$$E^{-1} = E_0^{-1} \left(1 + \frac{2E_0}{m_N} \sin^2 \frac{\theta}{2} \right)$$

Nuclear Form factor:

$$M_{fi}^{(0)} = \int \psi_f^* e \phi(\vec{x}) \psi_i d^3x$$

$$\nabla^2 \phi(\vec{x}) = -4\pi z e f(\vec{x}) ; \int_0^\infty f(\vec{x}) d^3x = 1$$

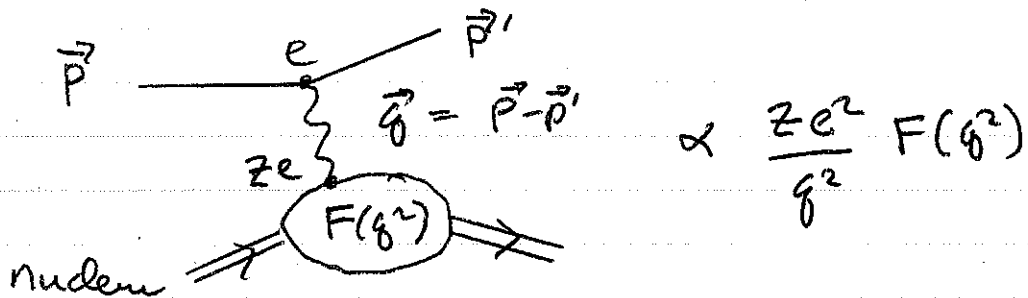
Use Green's theorem: $\int (u \nabla^2 v - v \nabla^2 u) d^3x = 0$

to get

$$M_{fi}^{(0)} = \frac{-4\pi z e^2}{V} \frac{\hbar^2}{q^2} \int e^{i\vec{q}\cdot\vec{x}/\hbar} f(x) d^3x$$

nuclear form factor $\equiv F(q^2)$

as Feynman-like graph, point charge $\rightarrow F(q^2)$



$$\frac{d\sigma}{d\Omega} = \frac{d\sigma}{d\Omega}^{\text{mott}} |F(q^2)|^2$$

For $f(x) = f_0 e^{-x/a}$

$$F(q^2) = \left(1 + \frac{-q^2}{a^2}\right)^{-1} \quad \text{"dipole"}$$

better approximation,

$$f(x) = f_0 \left[1 + \exp\left(\frac{r-c}{a}\right)\right]^{-1}$$

data given $c \approx 1.07 \text{ fm } A^{1/3}$; $a = 0.54 \text{ fm}$

see figure on next page.

Homogeneous charged sphere approximation,

$$r_{\text{rms}} = \sqrt{\langle r^2 \rangle} = 1.2 \text{ fm } A^{1/3}$$

from Povh, Rith, Scholz, Zetsche

68 5 Geometric Shapes of Nuclei

- Nuclei are not spheres with a sharply defined surface. In their interior, the charge density is nearly constant. At the surface the charge density falls off over a relatively large range. The radial charge distribution can be described to good approximation by a Fermi function with two parameters

$$\rho(r) = \frac{\rho(0)}{1 + e^{(r-c)/a}} \quad (5.52)$$

This is shown in Fig. 5.8 for different nuclei.

- The constant c is the radius at which $\rho(r)$ has decreased by one half. Empirically, for larger nuclei, c and a are measured to be:

$$c = 1.07 \text{ fm} \cdot A^{1/3}, \quad a = 0.54 \text{ fm}. \quad (5.53)$$

- From this charge density, the mean square radius can be calculated. Approximately, for medium and heavy nuclei:

$$\langle r^2 \rangle^{1/2} = r_0 \cdot A^{1/3} \quad \text{where } r_0 = 0.94 \text{ fm}. \quad (5.54)$$

The nucleus is often approximated by a homogeneously charged sphere. The radius R of this sphere is then quoted as the nuclear radius. The following connection exists between this radius and the mean square radius:

$$R^2 = \frac{5}{3} \langle r^2 \rangle. \quad (5.55)$$

Quantitatively we have:

$$R = 1.21 \cdot A^{1/3} \text{ fm}. \quad (5.56)$$

This definition of the radius is used in the mass formula (2.8).

- The surface thickness t is defined as the thickness of the layer over which the charge density drops from 90% to 10% of its maximal value:

$$t = r(\rho/\rho_0=0.1) - r(\rho/\rho_0=0.9). \quad (5.57)$$

Its value is roughly the same for all heavy nuclei, namely:

$$t = 2a \cdot \ln 9 \approx 2.40 \text{ fm}. \quad (5.58)$$

