

Lecture 6: Nuclear Physics

Fundamental physics Quantum Chromodynamics (QCD) highly nonlinear theory that cannot be solved to yield potential.

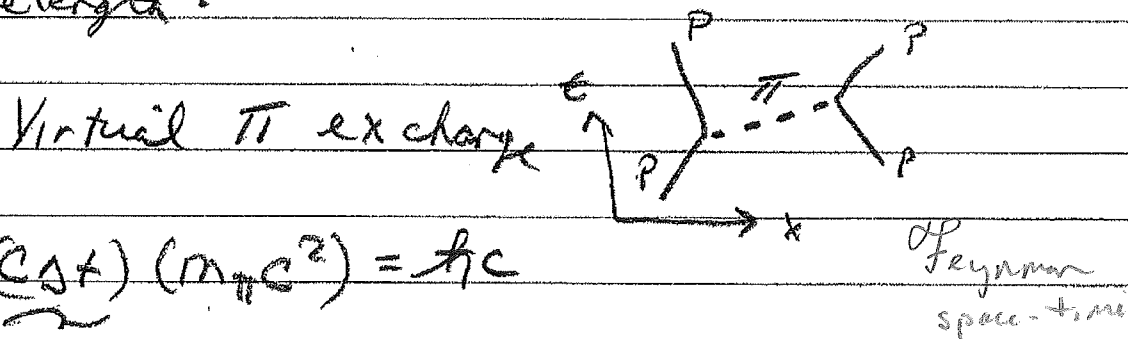
Approximate models (effective theories) based on meson (strongly interacting boson) exchange yield good description.

Basic properties of strong-nuclear force:

① Short range

$$V(r) = \frac{g^2}{r} e^{-r(m_\pi c/\hbar)}$$

range set by π -meson (pion) Compton wavelength.



$$\underbrace{(\hbar c)}_{\text{range}} (m_\pi c^2) = \hbar c$$

range set by uncertainty principle.

$$\frac{\hbar c}{m_\pi c^2} = \frac{200 \text{ MeV} \cdot \text{fm}}{140 \text{ MeV}} \approx \text{fm}$$

② Strong Coupling

$$\alpha_s = \frac{g^2}{4\pi\hbar c} \approx 15 \gg \alpha_{Em} = \frac{1}{137}$$

Expect binding energies $\sim \alpha_s^2 = (2 \times 10^3)^2 \alpha_{Em}$

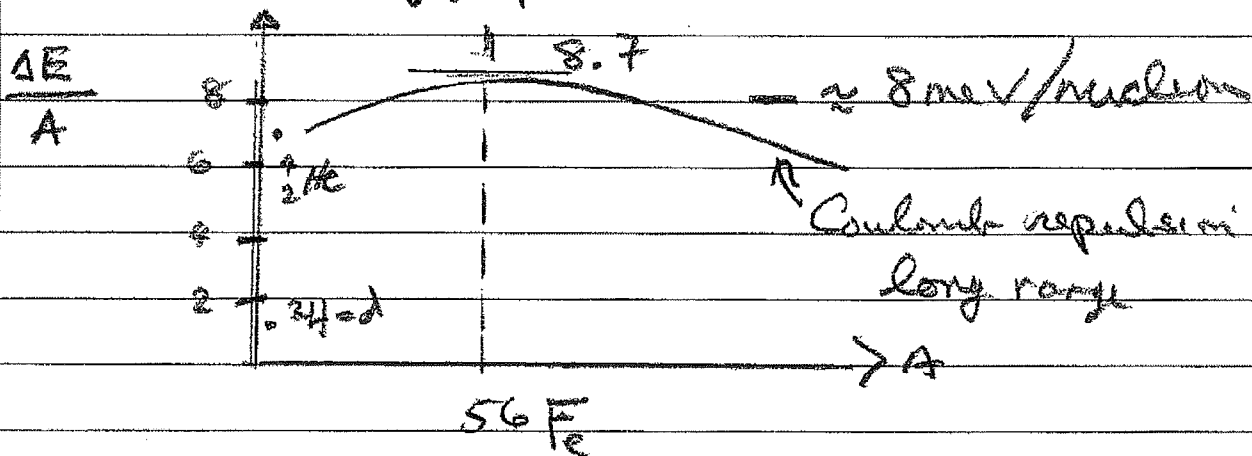
Order of $\sim 10^6 \times (10 \text{ eV atomic}) = \underline{10 \text{ MeV}}$,
magnitude

③ Nuclear radius

$$r_A = (1.2 \text{ fm}) A^{1/3} \quad \text{like dense sphere packing}$$

nuclear density $\approx 10^{15} \text{ g/cm}^3$

④ Binding energy per nucleon

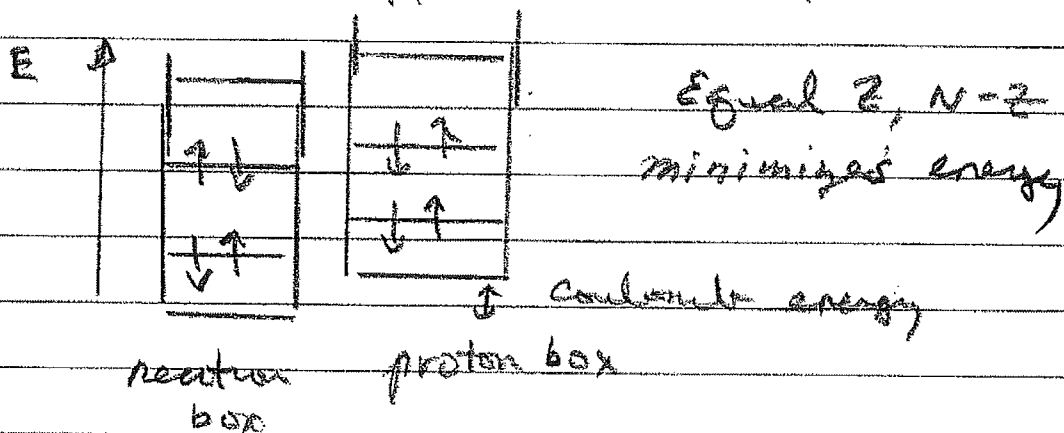


(i) Saturation due to short range

(ii) long range Coulomb \Rightarrow more neutrons needed to build heavy nuclei

⑤ Charge independence

light nuclei have $Z \approx (A-Z)$.
 result of Pauli exclusion principle
 and "charge independence" of nuclear
 forces verified in
 nn, np, pp scattering



"charge independence" can be implemented
 formally as isospin symmetry

$$N = \begin{pmatrix} p \\ n \end{pmatrix} \quad I = \frac{1}{2}$$

$$\pi = \begin{pmatrix} \pi^+ \\ \pi^0 \\ \pi^- \end{pmatrix} \quad I = 1$$

Write nuclear Hamiltonian in isospin
 invariant form. State must be
 completely antisymmetric wave function:

$$\Psi = (\text{space}) (\text{spin}) (\text{isospin})$$

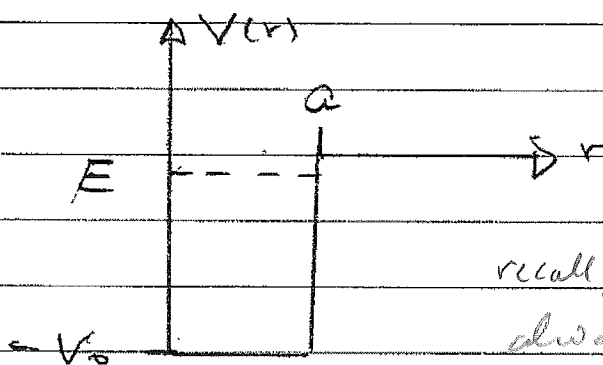
Finite Spherical Well model of deuteron

$d = np$ very weakly bound (few meV)
 only ground state ($l=0$) exists

$$\psi_0 = \frac{U}{r} \frac{1}{\sqrt{4\pi}}$$

∞ Finite spherical well potential model:

$$V(r) = \begin{cases} -V_0 & , r < a \\ 0 & , r > a \end{cases}$$



recall, in 1D bound state
 always exists. Not so
 in 3D

radial equation for $U(r)$ ($l=0$)

$$-\frac{\hbar^2}{2\mu} \frac{d^2 U}{dr^2} - V(r) U(r) = E U(r)$$

define $q = \sqrt{2\mu |E|} / \hbar$ $r > a$

$$Q = \sqrt{2\mu (V_0 - |E|)} / \hbar$$
 $r < a$

general solution

$$U(r) = \begin{cases} A \sin Qr + B \cos Qr & r < a \\ C e^{-Qr} + D e^{+Qr} & r > a \end{cases}$$

0 to satisfy $U(0) = 0$
boundary

0 normalizable

just as for Coulomb case, $U \rightarrow r^{l+1}$
 $r \rightarrow 0$

Suppose $U(0) = \text{constant}$.

then $\psi \rightarrow \frac{\text{const}}{r}$
 $r \rightarrow 0$

But $\boxed{\nabla^2 \left(\frac{1}{r}\right) = -4\pi \delta^3(\vec{r})}$ which does

not solve finite well differential equation.
(or Coulomb potential for that matter)

proof: divergence theorem

$$\int_V d^3r \nabla^2 \left(\frac{1}{r}\right) = \oint_S d^2A \cdot \hat{n} \cdot \vec{\nabla} \left(\frac{1}{r}\right)$$

choose V as a sphere of radius r :

$$\begin{aligned} \oint d^2A \hat{n} \cdot \vec{\nabla} \left(\frac{1}{r}\right) &= \int d^2A \frac{d}{dr} \left(\frac{1}{r}\right) = - \int \frac{d^2A}{r^2} = -4\pi \\ &= -4\pi \int d^3r \delta^3(\vec{r}) \end{aligned}$$

transcendental equation

match boundary conditions

$$A \sin Qa = C e^{-ga}$$

$$AQ \cos Qa = -g C e^{-ga}$$

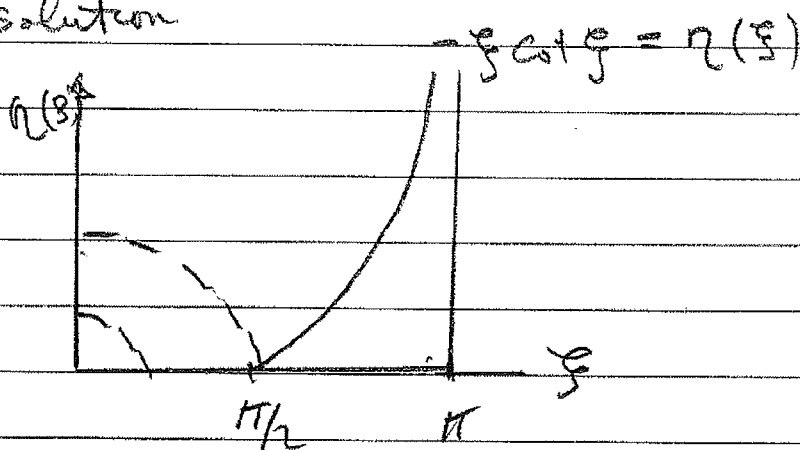
$$aQ \cot(Qa) = -g a$$

define $\eta = g a$, $\xi = Q a$

$$\xi \cot \xi = -\eta$$

$$\xi^2 + \eta^2 = \frac{2\mu}{\hbar^2} V_0 a^2 > \left(\frac{\pi}{2}\right)^2$$

graphical solution:



no bound state for $\xi^2 + \eta^2 < \left(\frac{\pi}{2}\right)^2$

in contrast to 1D square well

Estimate of V_0

photo-disintegration of d gives $E_\gamma = 2.2 \text{ MeV}$



So
$$\frac{2\mu}{\hbar^2} V_0 a^2 \approx \left(\frac{\pi}{2}\right)^2$$

exp. value of $a = 1.7 \text{ fm}$

$$V_0 a^2 = \left(\frac{\pi}{2}\right)^2 \frac{(197 \text{ MeV}\cdot\text{fm})^2}{938 \text{ MeV}} \approx 100 \text{ MeV}\cdot\text{fm}^2$$

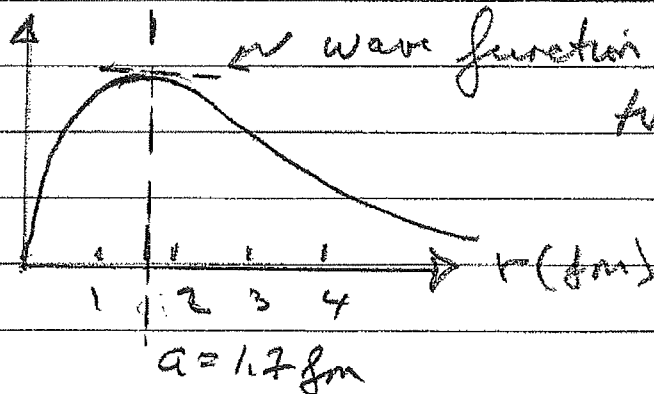
$$V_0 = 35 \text{ MeV}$$

$$Q^{-1} = \frac{\hbar c}{\sqrt{2\mu c^2 (V_0 - |E|)}} = \frac{200 \text{ MeV}\cdot\text{fm}}{\sqrt{938 \text{ MeV} \cdot (35 - 2) \text{ MeV}}}$$

$$= 1.1 \text{ fm}$$

$$q^{-1} = \frac{\hbar c}{\sqrt{938 \text{ MeV} (2.2 \text{ MeV})}} = 4.4 \text{ fm}$$

$U(r)$



more realistic nuclear force

Nuclear force is "charge-independent" isospin conserving, but spin dependent.

meson-theoretic two nucleon potential, one pion exchange in static limit.

Spin, isospin operators:

σ_i - Pauli matrices acting on spin state

τ_i - Pauli matrices acting on isospin state

dimensionless separation variable $X = r \left(\frac{m_\pi c^2}{\hbar c} \right) \sim 1/\text{fm}$

Spin tensor operator:

$$\hat{S}_{12} = 3(\vec{\sigma}_1 \cdot \hat{e}_r)(\vec{\sigma}_2 \cdot \hat{e}_r) - \vec{\sigma}_1 \cdot \vec{\sigma}_2$$

then (See for example Roy, Nigam Nuclear Physics, Wiley)

$$V(x) = \frac{1}{3} \left(\frac{g^2}{4\pi} \right) m_\pi c^2 \underbrace{\vec{\tau}_1 \cdot \vec{\tau}_2}_{\text{constant} \equiv A} \left\{ \vec{\sigma}_1 \cdot \vec{\sigma}_2 + \hat{S}_{12} \left(1 + \frac{3}{x} + \frac{3}{x^2} \right) \right\} \frac{e^{-x}}{x}$$

Wave function of 2 nucleon system

$$\Psi_{NN} = \Psi_{\text{space}} \chi_{\text{spin}} \lambda_{\text{isospin}}$$

In isospin formalism proton is isospin \uparrow , neutron isospin \downarrow

NN treated as identical particles.

Pauli-exclusion principle $\Psi_{NN}(1,2) = -\Psi_{NN}(2,1)$
completely anti-symmetric.

Expect ground state to be $l=0$ symmetric.

Aside: Actually deuteron has small $l=2$ component, giving it a small electric quadrupole moment, $Q \approx 0.3 e \cdot \text{fm}^2$

Iso spin $I=0$ odd $\chi_0^I = \frac{1}{\sqrt{2}} (\uparrow\downarrow - \downarrow\uparrow)$ (deuteron d)

$I=1$ even $\chi_{1,1}^I = \uparrow\uparrow$

$\chi_{1,0}^I = \frac{1}{\sqrt{2}} (\uparrow\downarrow + \downarrow\uparrow)$

$\chi^I = \downarrow\downarrow$ di-neutron

deuteron wave function: $l=0$

$\psi_d \approx \frac{u(r)}{r} \frac{1}{\sqrt{00}} \chi_0^I \chi_{1,m}^S \quad I=0, S=1$

di-neutron wave function:

$\psi_{nn} = \frac{u(r)}{r} \frac{1}{\sqrt{00}} \chi_{1,-1}^I \chi_0^S \quad I=1, S=0$

recall $2\hat{S}_1 \cdot \hat{S}_2 = \hat{S}^2 - \hat{S}_1^2 - \hat{S}_2^2 \quad \hat{S} = \hat{S}_1 + \hat{S}_2$

with $\hat{S}_i = \frac{\hbar}{2} \vec{\sigma}_i$

$2 \langle \frac{1}{2} \vec{\sigma}_1 \cdot \frac{1}{2} \vec{\sigma}_2 \rangle = S(S+1) - 3/2$

$\langle \vec{\sigma}_1 \cdot \vec{\sigma}_2 \rangle = 2 [S(S+1) - 3/2] = \begin{cases} -3 & S=0 \\ +1 & S=1 \end{cases}$

similarly for $\langle \vec{T}_1 \cdot \vec{T}_2 \rangle$

tensor operator \hat{S}_{12} :

choose spin-quantization direction as \hat{e}_r

S=0 singlet case

$$\langle \hat{S}_{12} \rangle = 3 \underbrace{\langle \vec{\sigma}_1 \cdot \hat{e}_r \vec{\sigma}_2 \cdot \hat{e}_r \rangle}_{-1} - \underbrace{\langle \vec{\sigma}_1 \cdot \vec{\sigma}_2 \rangle}_{-3} = 0$$

S=1 triplet case

$$\chi_{1, \pm 1}^S \quad \langle \hat{S}_{12} \rangle = 3 \underbrace{\langle \vec{\sigma}_1 \cdot \hat{e}_r \vec{\sigma}_2 \cdot \hat{e}_r \rangle}_{+1} - \underbrace{\langle \vec{\sigma}_1 \cdot \vec{\sigma}_2 \rangle}_{+1} = 2$$

$$\chi_{1, 0}^S \quad \langle \hat{S}_{12} \rangle = 3 \underbrace{\langle \vec{\sigma}_1 \cdot \hat{e}_r \vec{\sigma}_2 \cdot \hat{e}_r \rangle}_{-1} - \underbrace{\langle \vec{\sigma}_1 \cdot \vec{\sigma}_2 \rangle}_{+1} = -4$$

then deuteron bound state $(\chi_{1, \pm 1}^S) \quad I=0, S=1$

$$V_d(x) = A(-3) \left\{ 1 + 2 \left(1 + \frac{3}{x} + \frac{3}{x^2} \right) \right\} \frac{e^{-x}}{x}$$

$$= A(-3) \left\{ 3 + \frac{6}{x} + \frac{6}{x^2} \right\} \frac{e^{-x}}{x}$$

di-neutron state $I=1, S=0$

$$V_{nn}(x) = A(+1) \left\{ -3 \right\} \frac{e^{-x}}{x}$$

$V_d \sim 3 V_{nn}$ since d is weakly bound
 $n+d \rightarrow p+n$ threshold 2.2 MeV

\Rightarrow di-neutron is not bound.