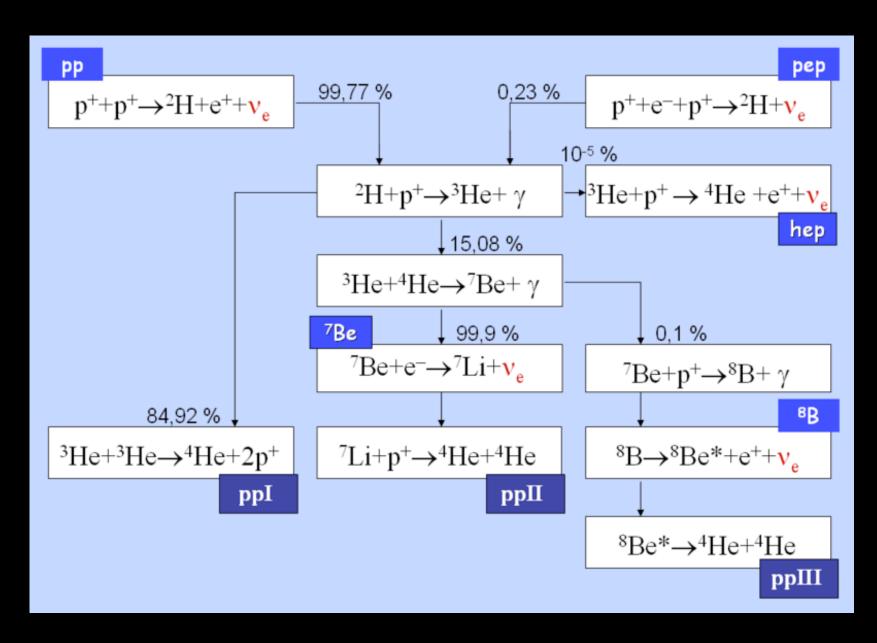
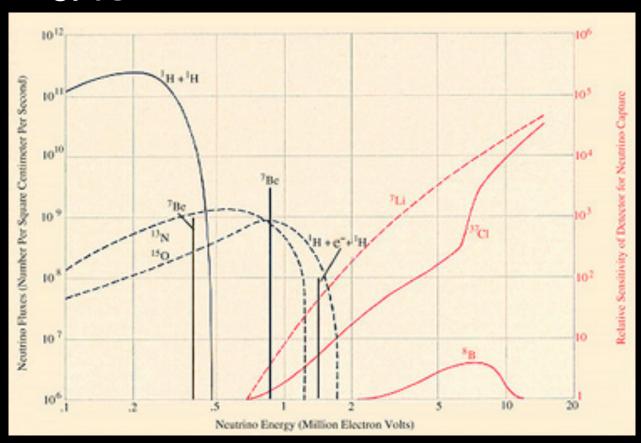
The Solar Neutrino Problem

effective solar fusion reaction
$$4p \rightarrow {}^{4}He + 2e^{+} 2v_{e} + 25 MeV$$



$SNU = 10^{-36} \text{ events/atom/s}$



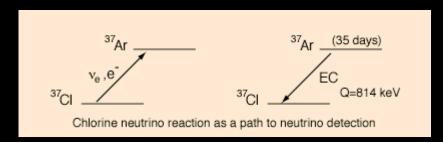
SPECTRUM OF SOLAR-NEUTRINO ENERGIES is plotted with curves showing the sensitivity of the ³⁷Cl detection system now in use (solid line in color) and the sensitivity of a proposed detection system employing lithium, ⁷Li (broken line in color). Neither system is sensitive in the region below about .8 MeV, where the energies of most of the solar neutrinos would be found. The lithium system, however, would be more sensitive than the ³⁷Cl system to neutrinos produced by the pep reaction, ¹H + e⁻ + ¹H. Most of the neutrinos expected to be captured by ³⁷Cl are those released by the decay of ⁸B. Neutrinos from the proton-proton chain are indicated by solid black lines, neutrinos from the CNO cycle by broken lines. The neutrino fluxes are plotted as the number per square centimeter per second per MeV for continuum sources and as the number per square centimeter per second for line sources.

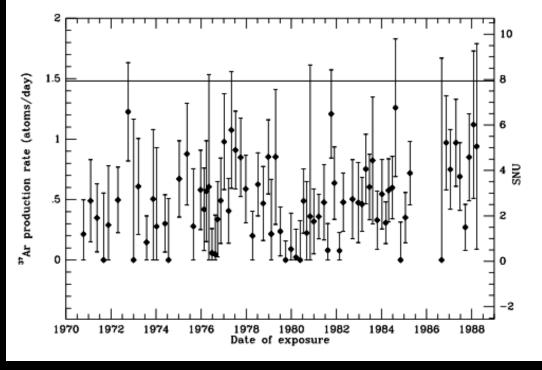
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1964

Raymond Davis of Brookhaven National Laboratory constructed a <u>neutrino</u> detector 1.6 km underground in the Homestake Gold Mine in Lead, South Dakota. The detector consists of a 378,000 liter tank of perchloroethylene, which is further isolated by being submerged in water. Theoretical expections were about one neutrino-chlorine interaction per day, but the measured solar neutrino events were about a third of that, raising serious questions about the abundance of solar neutrinos





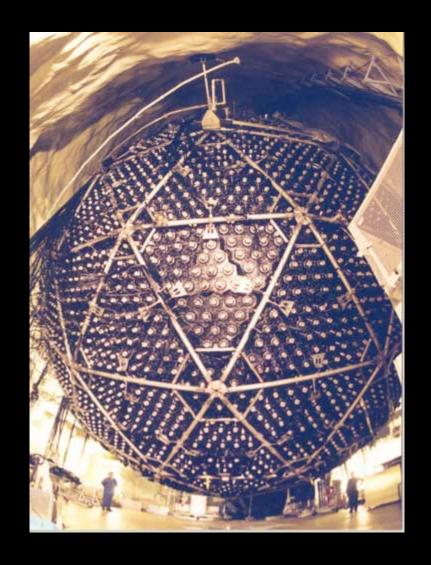


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$$v_e \rightarrow W^+ + e^ v_\mu \rightarrow W^+ + \mu^ v_\tau \rightarrow W^+ + \tau^ v \rightarrow Z + v$$

Sudbury Neutrino Observatory, a 12-meter sphere filled with heavy water surrounded by light detectors located two thousand meters below the ground in Sudbury, Ontario, Canada.

$$d+v\rightarrow p+n+v$$

Total Rates: Standard Model vs. Experiment Bahcall-Serenelli 2005 [BS05(OP)]

